

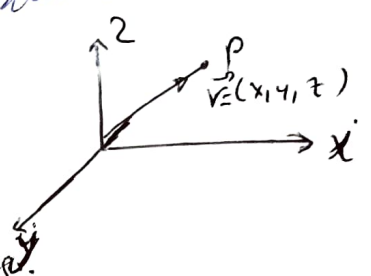
$\int \int \int \dots$

Kinetic model (Statistical mechanics) mesoscopic

* Phase space:

i- Configuration space: At any instant of time, each particle in the plasma can be localized by a position vector (\vec{r}) drawn from the origin of a coordinate system to the center of mass of the particle. In Cartesian coordinates:

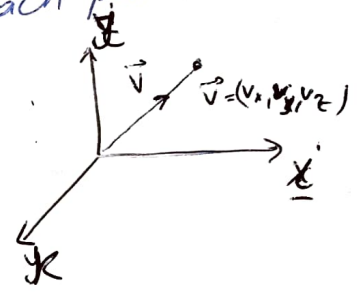
$$\vec{r} = x \underline{i} + y \underline{j} + z \underline{k}$$



- No of dimensions: $3N$; N : no of particles.

ii- Velocity space: At any instant of time, each particle can be localized by a velocity vector (\vec{v}):

$$\vec{v} = v_x \underline{i} + v_y \underline{j} + v_z \underline{k}$$



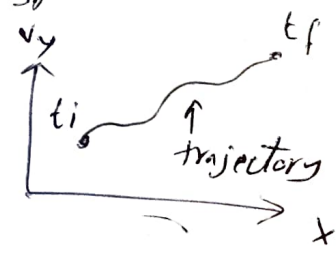
- No of dimensions: $3N$; N : no of particles.

3- Phase space:

- The instantaneous dynamic state of each particle can be specified by its position and velocity vectors, so the phase

Phase space = Configuration space + velocity space

- For each particle you need = 6D.
- For N particles = $6ND$.



- No of dimensions: $6ND$.
- Each ~~particle~~ system represents as a point, when the system moves, it forms a trajectory in phase space

Note that:

- (i) N space: 6D & each particle represent a point & for $N \rightarrow N$ points.
- (ii) N space: 6ND & each system \approx \approx for $N \rightarrow 1$ point.

* Differential elements:

~~The conf:~~

1- Configuration volume element:

- A small element of volume in configuration space is represented by $d^3r = dx dy dz$.

- This differential element d^3r is

- a mathematically infinitesimal quantity.
- finite element, sufficiently large to

contain a very large number of particles.

- Sufficiently small in comparison with the spatial variation of physical quantity, i.e. the physical doesn't change inside the element.

• Exp: In gas containing 10^{18} molecules, if we take $d^3r = 10^{-12} m^3$ there are 10^6 molecules inside $d^3r \rightarrow$ 10^{12} ensemble point around a terminal point

→ A Particle at Position \vec{r} inside d^3r , means

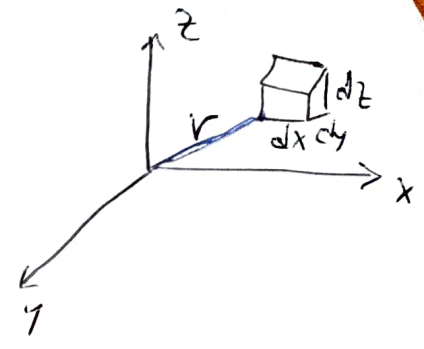
$$\vec{r} = x\hat{i} + y\hat{j} + z\hat{k}$$

no. of particles inside d^3r :
 $n = 10 \frac{\#}{m^3}$

- the x coordinate of the particle lies between x & $x + dx$.
- " y " " " " " " " y , $y + dy$.
- " z " " " " " " " " z , $z + dz$.

$$N = 10^{18} \#$$

the particle has arbitrary velocity in velocity space. = 10^{12} no. of ensembles



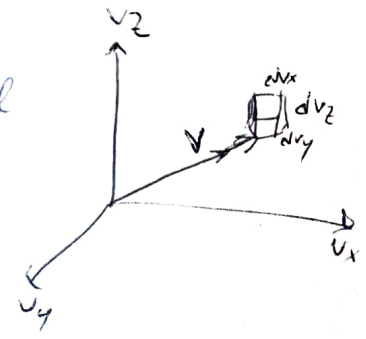
2- Velocity volume element:

- A small element of volume in velocity space is represented by

$$d^3v = dv_x dv_y dv_z$$

- A Particle inside d^3v , around the terminal point of the velocity vector \vec{v} .

- its v_x velocity lies between v_x and $v_x + dv_x$
- " v_y " " " " " " v_y and $v_y + dv_y$
- " v_z " " " " " " v_z and $v_z + dv_z$.



Phase Volume element:

In phase space (μ-space) a differential element of volume is a six-dimensional cube:

$$dV = d^3r d^3v = dx dy dz dv_x dv_y dv_z$$

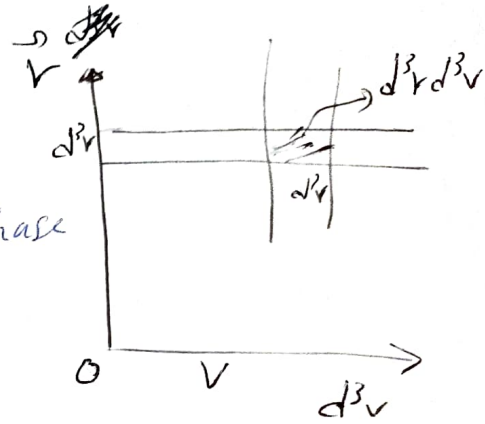
4 - differential number: ~~density~~ dN

- the number of particles inside the volume element $d^3r d^3v$ around the phase space coordinates (\vec{r}, \vec{v}) at instant t is

$$d^6 N(\vec{r}, \vec{v}, t)$$

- the total number is

$$N = \int d^6 N d^3r d^3v$$



* Distribution Function:

- the phase space ~~density~~ distribution function is given by

$$f(\vec{r}, \vec{v}, t) = \frac{d^6 N(\vec{r}, \vec{v}, t)}{d^3r d^3v} = \frac{\text{no. of particles}}{\text{Volume element}} \quad \boxed{d^6 N} \quad d^3r d^3v$$

- where: $f(\vec{r}, \vec{v}, t)$: the number density of particles inside $d^3r d^3v$.

- mass density:

$$\rho = \frac{M}{V}$$

$$dm = \rho dv$$

• if the density is not homogenous

$$dm = \rho(x) dv$$

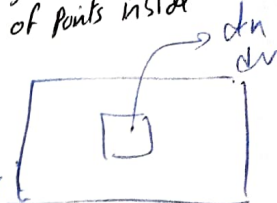


$f(\vec{r}, \vec{v}, t)$	$n(\vec{r}, t)$
• Number density function in phase space	• Configuration space number density
• $\frac{\#}{m^3 v^3} = \frac{N^6}{V^6}$	• $\frac{\#}{m^3} = \frac{N^3}{V^3}$
• distributed	• Random function.

- Number density:

$$f = \frac{d^6 N}{d^3 r d^3 v} = \frac{d^6 N}{d^6 \Gamma} \approx \frac{N}{V}$$

$d^6 \Gamma$: have to be small compared to the overall spatial extension of the points, but still keep it large enough to have a sufficient large number of points inside the volume



$d^6 \Gamma$: Volume of the ensemble in phase space has 6 D.

- $d^6 \Gamma = f d^6 \Gamma \rightarrow$ constant

- if the configuration space is not homogeneous

$d^6 \Gamma = f(x) d^6 \Gamma \rightarrow$ function of x .

- if velocity space is not isotropic:

$d^6 \Gamma = f(\vec{r}, \vec{v}) d^6 \Gamma$; function of \vec{r}, \vec{v}

$$f(\vec{r}, \vec{v}, t) = \lim_{\delta V \rightarrow 0} \frac{\delta^6 N}{\delta V}$$

- if the density is not static,

$d^6 \Gamma = f(\vec{r}, \vec{v}, t) d^6 \Gamma$; function of \vec{r}, \vec{v}, t .

So, $f(\vec{r}, \vec{v}, t)$: phase space density function

- if f is distributed in specific way, so $f(\vec{r}, \vec{v}, t)$ is called: phase-space density distribution function

- probability distribution function
total number of particles in phase space

$$N = \iint d^6 N = \iint f(\vec{r}, \vec{v}, t) d^3 r d^3 v$$

$$1 = \iint \frac{f(\vec{r}, \vec{v}, t)}{N} d^3 r d^3 v$$

$$= \iint P(\vec{r}, \vec{v}, t) d^3 r d^3 v$$

$n(\vec{r}, t) = \int f(\vec{r}, \vec{v}, t) d^3 v$
↳ configuration space number density function

$P(\vec{r}, \vec{v}, t)$: Probability distribution function.

$$dP(\vec{r}+d\vec{r}, \vec{v}+d\vec{v}, t) = P(\vec{r}, \vec{v}, t) d^3 r d^3 v$$

Classification of distribution function:

Equilibrium vs Non equilibrium:

- If the distribution function f doesn't depend on t is called "Equilibrium distribution function".

- If " " " " f depend on t is called "Non equilibrium distribution function".

② Homogenous vs Non homogenous:

- If the distribution function f doesn't depend on \vec{r} , it is called "Homogenous distribution function".

- If the distribution " " f depends on \vec{r} , it is called "Non homogenous distribution function".

③ Isotropic vs Anisotropic:

- If the distribution function f depends only on ~~the~~ the magnitude of $v = |v|$, it is called "Isotropic distribution function".

- If " " " " depends on the orientation of \vec{v} , it is called "Anisotropic distribution function".

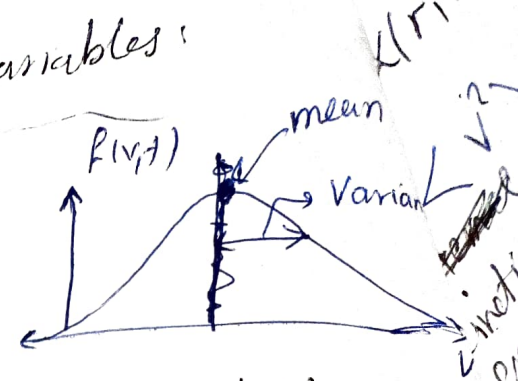
* Conditions of the distribution function:

- ① continuous
- ② real positive
- ③ finite.

* Moments of distribution function:
Average values and macroscopic variables:

(i) Moments:

$$M_n = \int v^n f(x, v, t) dv dx$$



- Zero moment $M_0 = \int v^0 f(x, v, t) dv = \text{Area under}$ (total number)
- 1st moment $M_1 = \int v^1 f(x, v, t) dv = \text{Average OR mean value}$ (momentum)
- 2nd moment $M_2 = \int v^2 f(x, v, t) dv = \text{Variance}$ (energy)
- 3rd moment $M_3 = \int v^3 f(x, v, t) dv = \text{Skewness}$ (Symmetry around mean).

(ii) Average values and macroscopic variables: (11 Pat)
 set of many replicas of the system, which are identical in all other aspects apart from being in different states at an instant of time.

- The average value of the property $\chi(\vec{r}, \vec{v}, t)$ for the particle

$$\langle \chi(\vec{r}, \vec{v}, t) \rangle = \frac{1}{n(\vec{r}, t)} \int \chi(\vec{r}, \vec{v}, t) f(\vec{r}, \vec{v}, t) d^3v$$

(i) $\chi(\vec{r}, \vec{v}, t) = v^0$

$$\langle v^0 \rangle = \frac{1}{n(\vec{r}, t)} \int v^0 f(\vec{r}, \vec{v}, t) d^3v = 1$$

$$\therefore n(\vec{r}, t) = \int f(\vec{r}, \vec{v}, t) d^3v = \frac{\text{number density. No of particles}}{d^3v} = \frac{\#}{m^3}$$

(ii) $\chi(\vec{r}, \vec{v}, t) = v^1$

$$\langle v^1 \rangle = \frac{1}{n(\vec{r}, t)} \int v^1 f(\vec{r}, \vec{v}, t) d^3v = \vec{u}$$

macroscopic average velocity OR flow velocity

$$\rho \vec{u} = m n(\vec{r}, t) \vec{u}(\vec{r}, t) = \int v^1 f(\vec{r}, \vec{v}, t) d^3v = \rho$$

momentum density

mean $\langle (\vec{r}, \vec{v}, t) \rangle = v^2 :$

varian $\langle v^2 \rangle = \frac{1}{n(\vec{r}, t)} \int v^2 f(\vec{r}, \vec{v}, t) d^3v = u^2$

kinetic energy density

$\frac{1}{2} m n(\vec{r}, t) u^2 = \int v^2 f(\vec{r}, \vec{v}, t) d^3v = \mathcal{R}$

* Note that: The peculiar velocity or random velocity \vec{c} :
the velocity of the particle relative to the average velocity:

$\vec{c} = \vec{v} - \vec{u} ; \langle c \rangle = 0 \text{ \& } \langle v \rangle = u$

* The pressure:

$\langle c^2 \rangle = \frac{1}{n(\vec{r}, t)} \int c^2 f(\vec{r}, \vec{v}, t) d^3v$

pressure

$\rho = n(\vec{r}, t) \langle c^2 \rangle$

* the temperature: is a measure of the mean kinetic energy of the random particle motion.

$\frac{1}{2} k_T = \frac{1}{2} m \langle c^2 \rangle$
 thermal energy \swarrow \searrow mean kinetic energy of random motion.

* Plasma kinetic equation:

- the plasma kinetic equation:

i- Compact form: $\frac{Df}{Dt} = C(f) = \left(\frac{df}{dt}\right)_{coll}$

ii- Regular form: $\frac{\partial f}{\partial t} + (\vec{v} \cdot \vec{\nabla}_r) f + (\vec{a} \cdot \vec{\nabla}_v) f = C(f)$

iii- Conservative form: $\frac{\partial f}{\partial t} + \underbrace{\vec{\nabla}_r \cdot (\vec{v} f)}_{\text{long-range interaction effect}} + \underbrace{\vec{\nabla}_v \cdot (\vec{a} f)}_{\text{short-range interaction effect}} = C(f)$

* where:

$\frac{D}{Dt} = \frac{\partial}{\partial t} + \vec{v} \cdot \vec{\nabla} + \vec{a} \cdot \vec{\nabla}$

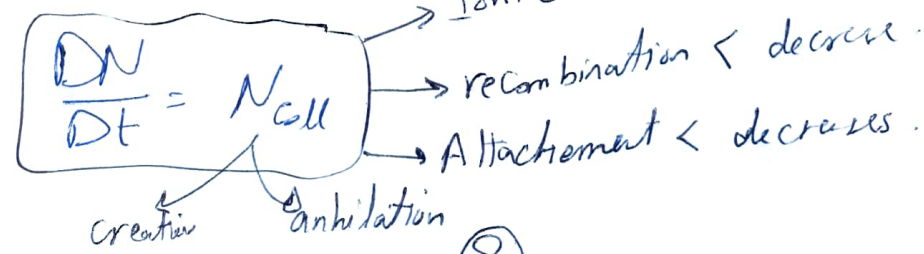
$C(f)$: Collision term $\begin{cases} \rightarrow \text{Boltzmann collision operator: Integral} \\ \rightarrow \text{Fokker-Planck } \dots : \text{differential} \\ \rightarrow \text{Krook } \dots : \text{Algebra} \end{cases}$

$\frac{\partial f}{\partial t}$: a Cumulation of f with time.

$(\vec{v} \cdot \vec{\nabla}_r)$: Advection of f with velocity v in configuration space
 : convection ~ ~ ~ ~ ~
 : Spatial acceleration

$(\vec{a} \cdot \vec{\nabla}_v)$: Advection of f with velocity a in velocity space.
 : convection ~ ~ ~ ~ ~
 : velocity acceleration.

$\frac{D}{Dt} \int f d^3r d^3v = \int C(f) d^3r d^3v$

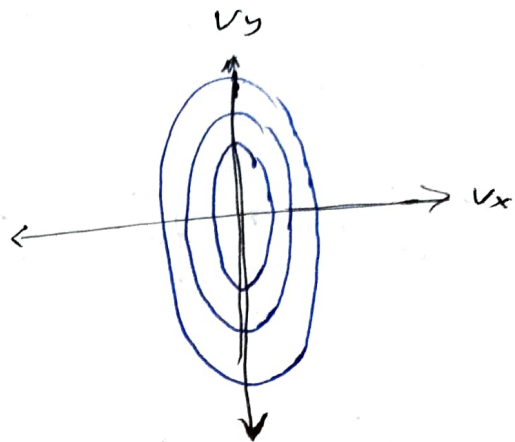
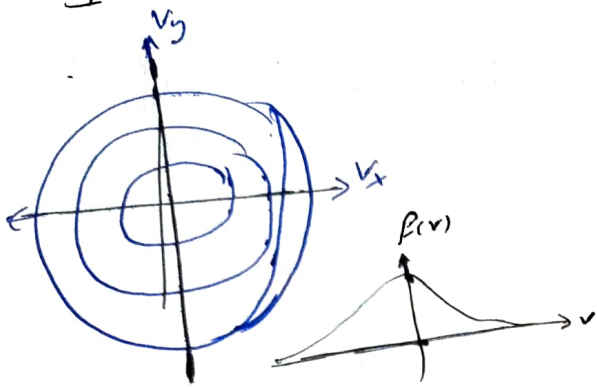


number

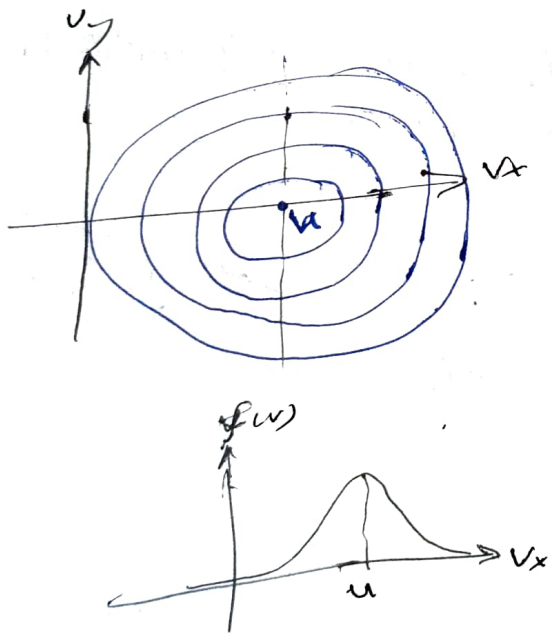
Types of distribution function:

① Anisotropic ($v_y > v_x$)

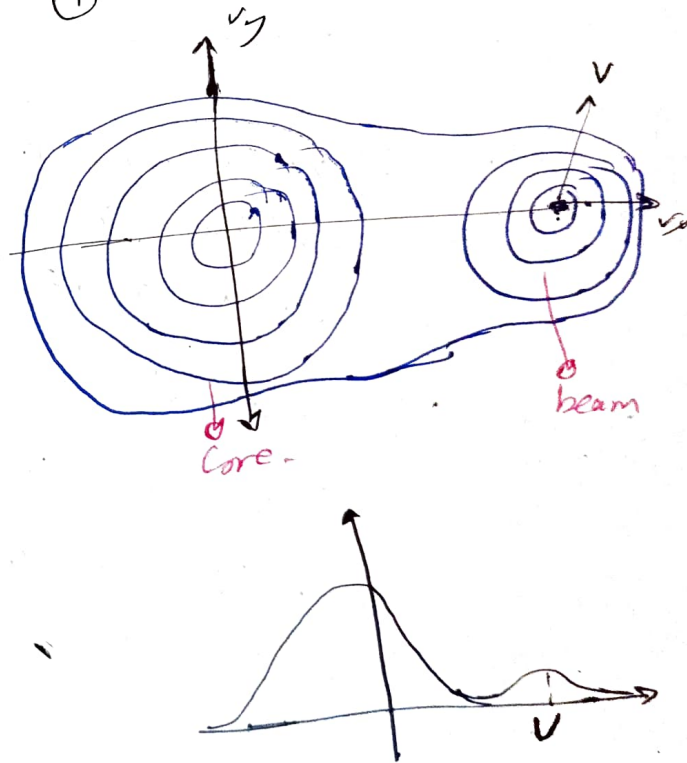
Isotropic Maxwellian



② Drift-Maxwellian



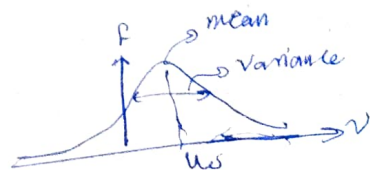
④ Bimodal / Beam



* Moments of distribution function

$$f(x, t) = C e^{-\frac{(u-u_0)^2}{\sigma^2}}$$

- Zero moment $\int f dv = C$: Area under
- 1st moment $\int u f dv = u_0$: Average or mean value or
- 2nd moment $\int u^2 f dv = \frac{1}{2} m u_0^2 + \frac{1}{2} P_0$ (variance)
- 3rd moment $\int v^3 f dv = 0$: Skewness (symmetry around mean)



* Plasma kinetic eq:

$$\frac{df}{dt} = \left(\frac{\partial f}{\partial t} \right)_c = C(f)$$

long-range interaction

- Boltzmann collision operator \rightarrow integral
- Fokker-Planck collision operator \rightarrow differential
- Krook collision operator: algebraic
- charge exchange collision operator \rightarrow integral

$$\frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} \frac{dx}{dt} + \frac{\partial f}{\partial v} \frac{dv}{dt} = C(f)$$

→ short-range interaction

$$\frac{\partial f}{\partial t} + (\underline{v} \cdot \underline{\nabla}_x) f + (\underline{a} \cdot \underline{\nabla}_v) f = C(f)$$

$\frac{\partial f}{\partial t}$: accumulation of f

$(\underline{v} \cdot \underline{\nabla}_x) f$: - advection of f with velocity v in configuration space
 - convection in configuration space.
 - spatial acceleration.

$(\underline{a} \cdot \underline{\nabla}_v) f$: - advection of f with velocity a in velocity space
 - convection term in velocity space.
 - velocity acceleration.

- It can be written as:

~~$$\frac{dN}{dt} = \iint C(f) d^3x d^3v$$~~

← creation or fusion or format
 ← annihilation or sink or loss

$$\frac{d}{dt} \iint d^3x d^3v f = \iint C(f) d^3x d^3v$$

$$\frac{df}{dt} = C(f)$$